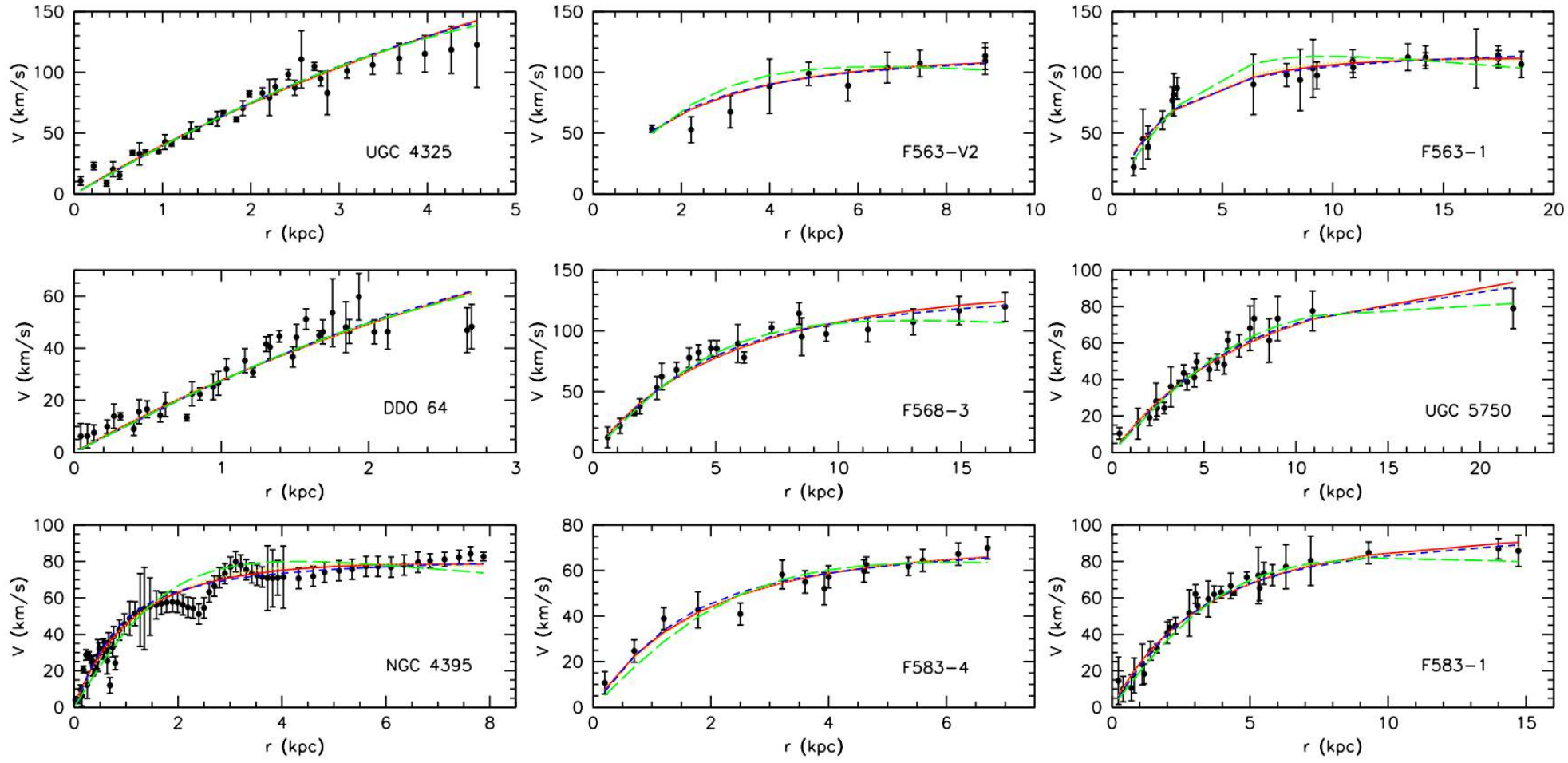


Dark matter cusps vs. cores

Observationally-inferred cores in low-surface brightness (LSB) galaxies



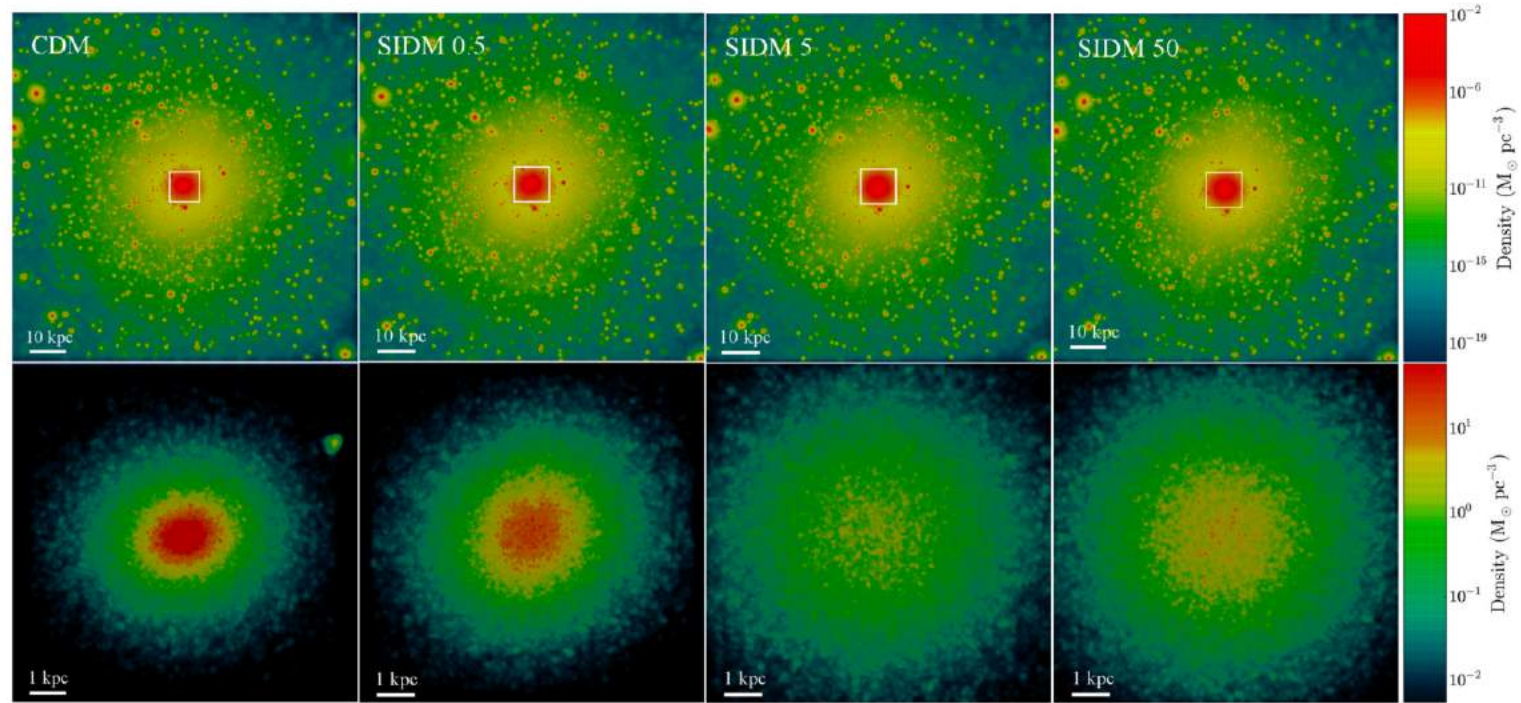
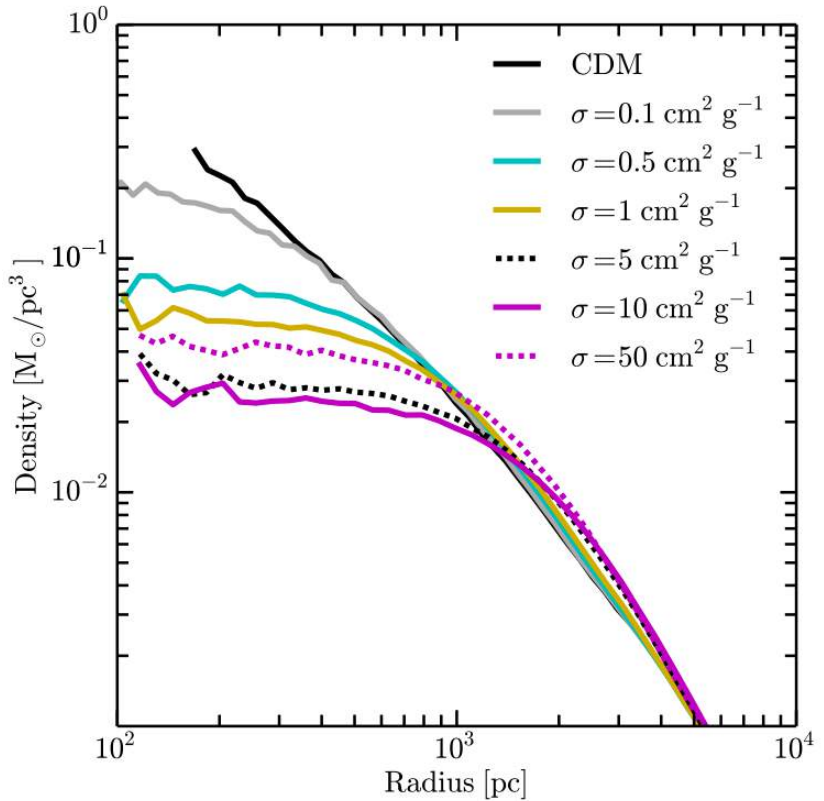
$v_c \propto r \Rightarrow \rho = \text{const}$

Compare with pure CDM: $\rho_{\text{NFW}} \propto r^{-1}$ for small r

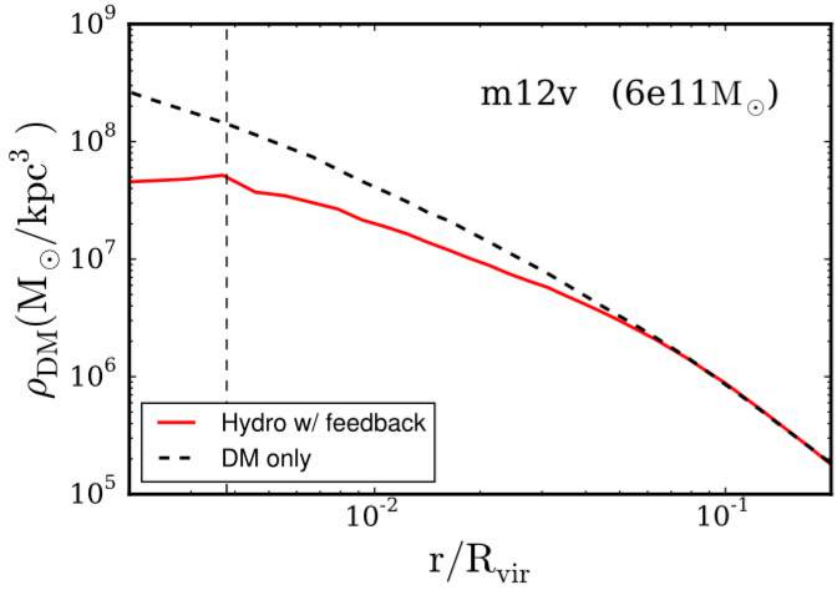
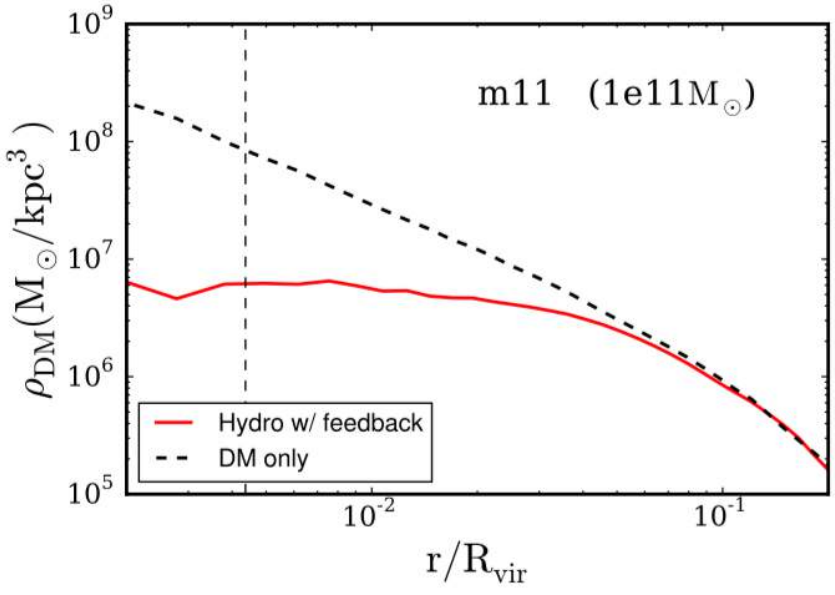
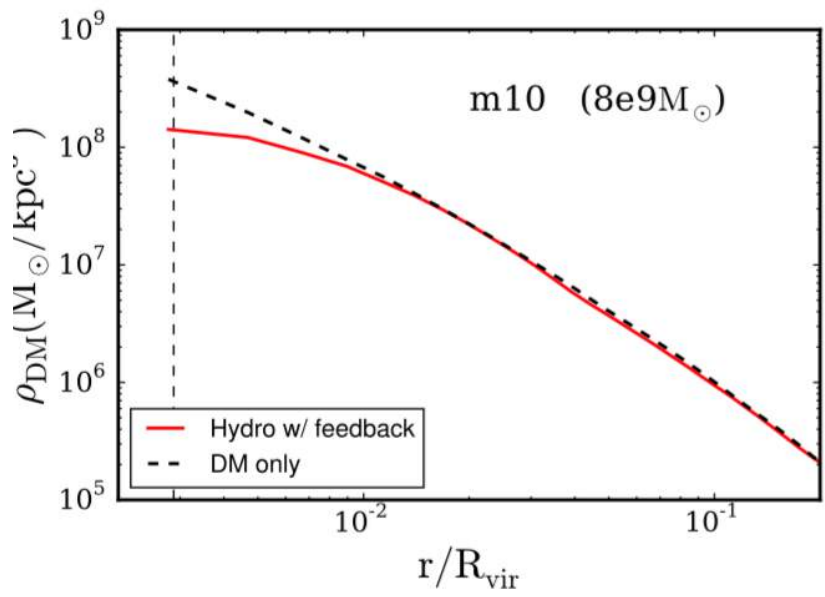
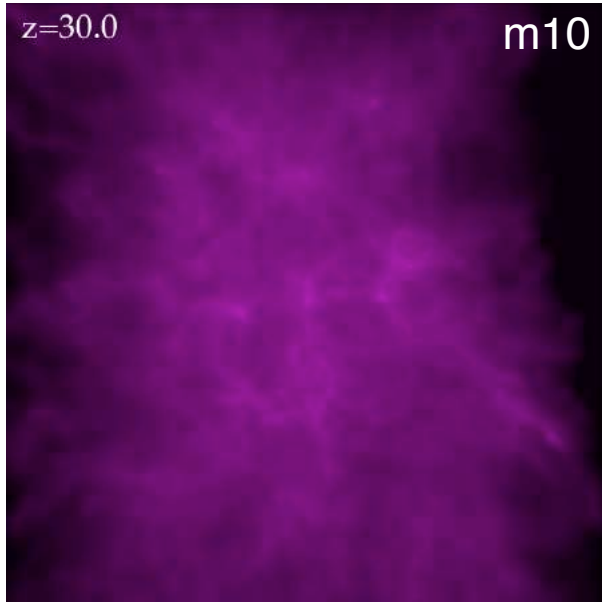
Self-interacting dark matter

One way to produce dark matter cores is to postulate that dark matter scatters with itself with a cross section σ (no annihilation nor dissipation needed)

Cons: correct σ unknown, values required to explain rotation curves not particularly well motivated



Stellar feedback can also produce cores by



Pros: appears to be an inevitable consequence of successful galaxy formation models, does not require “new physics”